

1 **Macro- and microphysical characteristics of rain cells observed during SOS-CHUVA**  
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18 **Key Points:**

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- Holistic view of convective rain cells in Campinas, Brazil using X-band radar.
  - The importance of the macroscale context for microphysical analyzes is explored.
  - Polarimetric and gamma DSD variables are studied in a Lagrangean way for rain cells.
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## 23 **Abstract**

24 In this study we present a methodology to study the properties of convective precipitation in a  
25 holistic way. We apply a tracking algorithm to X-band radar retrievals to store rain cells  
26 properties in a Lagrangean framework. The Center of Activity (COA), the altitude with the  
27 highest contribution to the cell's Vertically Integrated Liquid (VIL), is presented as a new  
28 perspective to study convective clouds. It is shown that the combination of COA and VIL  
29 provides a useful phase space to compare the properties of different rain cells. In our  
30 methodology, rain cells present high COA (around 4.0 km to 4.5 km) and high VIL ( $\sim 3.0 \text{ kg m}^{-2}$ )  
31 upon detection and evolve towards lower COA and VIL values throughout the lifecycle. The  
32 COA-VIL space is also used to constrain the microphysical study of the cells. Averaged  
33 polarimetric variables and gamma Droplet Size Distribution (DSD) parameters indicate that  
34 collection, melting and evaporation processes are almost in balance at COA independently of the  
35 rain cells lifecycle stage. High COA and negative DSD shape parameters ( $\mu$ ) at COA were found  
36 10 minutes before peak accumulated rainfall in low levels and are also likely associated to  
37 lightning activity and differential reflectivity ( $Z_{dr}$ ) columns. Overall, the results presented here  
38 can help nowcasting applications by providing expected microphysical characteristics from  
39 COA-VIL calculations. Contrary to microphysical retrievals, computations of COA and VIL do  
40 not depend on dual polarization.

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## 42 **Plain Language Summary**

43 In this study we looked for rain properties that are shared among different rain occasions during  
44 the summer of 2016/7 around Campinas city in Southeast Brazil. By following each case with a  
45 weather radar, we were able to store rain properties throughout their life cycle as well as in  
46 different altitudes. It was found that the altitude with the highest rain mass, together with the  
47 overall water mass amount, are two key properties that can help categorize rain occasions. When  
48 most of the water is relatively high in the clouds, e.g. 4 km high, the clouds were found to be in  
49 their most developed stage, where lightning activity is expected to start. About 10 minutes after  
50 this point, the bulk of the rain reaches lower levels (2 km in our methodology). This process is  
51 modulated by the overall rain water mass generated by the clouds, meaning that both properties  
52 should be used together for a more complete view. Additionally, this approach provides a context  
53 for more detailed studies of precipitation, where the processes responsible for rain formation are  
54 the main focus. We propose that the results presented here can help both operational forecasters  
55 and weather researchers to better anticipate intense rain occasions and the associated potential  
56 damages for society such as strong winds and lightning activity.

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## 62 **1 Introduction**

63 Convective clouds are one of the primary focus of meteorological studies given their  
64 importance to the overall climate and their effects on human society functioning. When severe  
65 convective clouds, either individually or in groups, act over populated areas, both human and  
66 patrimonial damage may follow. Perhaps the primary example are flood occasions, which have  
67 both direct and indirect effects – the first related to the elevation of rivers water levels and the  
68 latter to the dissemination of diseases by polluted rivers (Beyer, 1974). The progressive  
69 urbanization process, especially unorganized urbanization, intensifies the problem by limiting the  
70 surface infiltration capacity (Stevaux et al., 2009). Additionally, convective systems may also  
71 produce lightning activity, hail and windstorms and can disrupt cities electricity distribution.  
72 Hailfall, in particular, have a lot of damage reports in Brazil (Martins et al., 2017). Therefore,  
73 understanding the nature of convective systems is crucial in order to minimize their effects and  
74 improve the response time by providing early warnings.

75 Recent efforts to further the knowledge of convective clouds in Brazil were made around  
76 the deployment of the CHUVA project (the word for “rain” in Portuguese, an acronym for Cloud  
77 Processes of the Main Precipitation Systems in Brazil: A Contribution to Cloud-Resolving  
78 Modeling and to the Global Precipitation Measurement (GPM), Machado et al., 2014). Calheiros  
79 and Machado (2014) reported on the early results of the campaign, where precipitation clouds in  
80 different regions in Brazil were compared. They found that clouds over Southeast Brazil present  
81 the most active mixed layer, which may be due to enhanced pollution of the highly urbanized  
82 region. The CHUVA project lead to the implementation of a new experimental campaign  
83 focused primarily on Southeast Brazil named SOS-CHUVA (Nowcasting of intense  
84 thunderstorms and understanding of the physical processes inside clouds: the SOS-CHUVA  
85 (Severe Weather Observation System)). The overall aim of SOS-CHUVA is to study deep  
86 convective clouds around the region of Campinas (São Paulo state) using ground-based  
87 instrumentation including a X-band radar used in the present study.

88 Here we will report on X-band radar measurements during the summer of 2016/7 with the  
89 focus to highlight holistic characteristics of the precipitating clouds around Campinas. The  
90 progressive improvement of model resolutions and experimental capabilities have elucidated the  
91 highly complex nature of the macro- and microphysical properties of convective clouds. These  
92 have prompted the development of new strategies that aim to find key properties that synthesize  
93 the systems characteristics in a holistic way. An example is introduced in Heiblum et al. (2016a),  
94 where the authors use large eddy simulations to analyze warm cumulus fields in a 2-dimensional  
95 phase space defined by each cloud’s center of gravity (COG) altitude and their respective  
96 average liquid water path (LWP) or total amount of water mass. The authors highlight several  
97 cloud properties in the phase space, showing patterns typical of the growing or dissipating (both  
98 reversible and irreversible) stages, precipitating clouds, mergers and splits, for example.  
99 Therefore, it can be used as a tool to compare different clouds life cycles while interpreting the  
100 underlying physical and morphological aspects. One example of application of the COG-LWP  
101 phase space is provided in Heiblum et al. (2016b). The authors studied aerosol-cloud interactions  
102 in the phase space and showed that aerosols significantly alter the shapes and trajectories in it  
103 according to the underlying physical processes being affected. Overall, the COG tends to rise  
104 with aerosol loading in warm clouds because of both microscale processes such as decreased  
105 droplet terminal velocities and increased updraft speeds as well as macroscale effects in the  
106 increased environmental instability. The aerosol effect on the COG was further explored in Chen

107 et al. (2017), where the effects in the up- and downdrafts were quantitatively compared (e.g.  
108 increased droplet mobility and decreased terminal speed, respectively). Dagan et al. (2018)  
109 studied how the aerosol effect on the warm phase affected the mass transport to deeper systems  
110 that grow above the freezing level. They found that increased pollution tends to increase mass  
111 transport both up- and downwards through the freezing level as a result of the higher COG and  
112 the more active mixed phase. Preliminary SOS-CHUVA results indicate that aerosol  
113 concentrations revolve around a few thousands per  $\text{cm}^{-3}$ , which is consistent with urban centers  
114 environments. Therefore, it is expected that clouds analyzed here resemble the most polluted  
115 cases of the literature discussed above.

116 In terms of cloud microphysics, Cecchini et al. (2017) introduced the so-called Gamma  
117 phase space to provide a holistic way to analyze processes inside individual clouds. The authors  
118 fitted Gamma curves to aircraft in-situ measurements of cloud droplet size distributions (DSD)  
119 and used the three parameters - intercept, shape and curvature - to construct the phase space.  
120 Therefore, every DSD was viewed as a point in a 3D subspace and their evolution as trajectories.  
121 They argue that processes such as condensation and collision-coalescence produce specific  
122 patterns in the phase space, which could be used to both validate model calculations and to  
123 develop new parameterizations that also rely on the Gamma DSD approximation. In the present  
124 study, we will apply both concepts of macro- and microscale phase spaces to generate a  
125 framework adequate to holistically study the convective systems measured during SOS-CHUVA.  
126 Because the experimental approach is significantly different than the studies mentioned here,  
127 several practical and theoretical adaptations were necessary as discussed in the next sections.

128 Here we introduce the Center of Activity (COA) that is a measure of the clouds altitude  
129 level with the highest amount of water mass as seen from radar reflectivity measurements. This  
130 property, together with the Vertically Integrated Liquid (VIL) will be used as a macroscale  
131 constraint similar to the phase space introduced in Heiblum et al. (2016a). This approach  
132 provides a new perspective on the study of clouds lifecycle from radar measurements while  
133 providing a framework to analyze individual or groups of cases. Within this context,  
134 microphysical properties will be analyzed from the polarimetric radar measurements and Gamma  
135 DSD parameters estimates.

136 Section 2 provides information on data handling and analysis methodology overview.  
137 The results are shown in Section 3, while the conclusions are provided in Section 4.

## 138 **2 Methodology**

### 139 **2.1 Radar data**

140 The X-band dualpol radar was operated at 9.345 GHz with a  $1.3^\circ$  beamwidth at -3 dB and  
141 in a simultaneous transmission and reception mode, characteristic that allows to get the  
142 reflectivity at horizontal polarization (henceforth called  $Z$  for simplicity), differential reflectivity  
143  $Z_{dr}$ , the differential phase  $\phi_{dp}$ , among other variables. The gates resolution is 200 m, with a total  
144 covered range of 100 km. The raw radar dataset was processed to reduce the influence of the  
145 attenuation observed at X-band frequencies and to present better  $Z_{dr}$  estimates. To do so, the  
146 radar performed a zenith-pointed scan (also known as a “bird-bath” scan) after each volume  
147 scan. The  $Z_{dr}$  offset trend observed during time can be then corrected. To derive  $K_{dp}$  the  $\phi_{dp}$  was  
148 filtered and smoothed using a FIR (finite impulse response) filter and then the  $K_{dp}$  is estimated by

149 deriving the smoothed  $\varphi_{dp}$  curve (Hubbert & Bringi, 1995). Attenuation correction was applied to  
 150 both  $Z$  and  $Z_{dr}$  using the ZPHI method proposed by Testud et al. (2000).

151 From the processed radar data, we calculate  $Z$  CAPPIs (Constant Altitude Plan Position  
 152 Indicator) every 500 m, from 2 km up to 18.5 km. The CAPPIs have 1 km resolution, with the  
 153 radar at the center of the grid (22° 48' 50'' S, 47° 3' 22'' W). The volumetric data is interpolated  
 154 to the CAPPI grid by using a weighting function based on the distance between volumetric and  
 155 CAPPI pixels. In that way, the CAPPI grid points can be affected by multiple volumetric pixels,  
 156 with weight inversely proportional to their distance.

## 157 2.2 ForTraCC

158 In order to automatically track precipitating clouds, we employed the ForTraCC  
 159 (Forecasting and tracking Cloud Clusters – Vila et al., 2008) system. This algorithm was initially  
 160 developed to detect and track cloud clusters from GOES imagery but was later adapted to work  
 161 with radar CAPPIs by Queiroz (2009). In this study, we locate and track rain cells on CAPPIs at  
 162 2 km height (1 km horizontal resolution). ForTraCC is able to track rain cells by applying a  
 163 minimum threshold in rain rate ( $R$ ) or  $Z$ . Here we use the threshold of  $R = 5 \text{ mm h}^{-1}$  to identify  
 164 the cells, which is equivalent to around 34 dBZ according to the Marshall-Palmer  $Z$ - $R$  relation.  
 165 This is a relatively high threshold as compared to the more usual 20 dBZ limit. The intent is to  
 166 avoid too much chaotic mergers and splits and to reduce the noise in the overall statistics due to  
 167 high numbers of small rain cells. The result is that we effectively detect more mature rain cells  
 168 than usual, but their lifecycles are more clearly defined. Rain cells detected with our  
 169 methodology must present an area of at least 10 pixels (or  $10 \text{ km}^2$ ) to be tracked and mergers and  
 170 splits are tracked according to the maximum overlap area between sequential CAPPIs. Because  
 171 rain cells must conform to the threshold throughout the lifecycle, the 34 dBZ limit also means  
 172 the latter stages are not fully captured. However, the most intense stages of the cells are likely  
 173 better represented compared to a noisier 20 dBZ threshold.

174 CAPPIs are calculated every 10 minutes, enabling ForTraCC to store physical and  
 175 morphological characteristics of the rain cells throughout their lifecycle. Such characteristics  
 176 include cell's area evolution, movement speed and direction, among others. Therefore,  
 177 ForTraCC allows the definition of a Lagrangean dataset with rain cells characteristics. The  
 178 specific parameters obtained in this work are listed below.

179 Every rain cell identified is treated as a 3D cylinder, where its horizontal radius ( $r$ , in km)  
 180 is obtained from the cell area ( $A$ , in  $\text{km}^2$ ) provided by ForTraCC:

$$181 \quad r = \sqrt{\frac{A}{\pi}} + 2 \quad (1)$$

182 Where 2 km are added to  $r$  to compensate for possible vertical tilting of the clouds. By  
 183 using the same  $r$  in every altitude level, the cylinder is defined and it follows the cells throughout  
 184 their lifecycle.

185 The cylinders are used to obtain Lagrangean properties of the rain cells from the CAPPIs  
 186 as well as the volume scans. The properties calculated are separated between macro- and  
 187 microscale as listed below. Macroscopic characteristics obtained are:

- 188 a. Maximum area  $A_{max}$  (at 2 km), calculated as the maximum  $A$  in the cell lifecycle.

189 b. Duration  $\Delta T$  (in minutes), calculated by  $10(n - 1)$  where  $n$  is the number of  
190 CAPPIs in the cells lifecycle.

191 c. VIL (Vertically Integrated Liquid, in  $\text{kg m}^{-2}$ ), calculated by:

$$192 \quad VIL = \sum_{i=1}^{34} 3,44 \cdot 10^{-6} [(Z_i + Z_{i+1})/2]^{4/7} \cdot \Delta h \quad (2)$$

193 where  $i$  represents each of the 34 CAPPIs between 2 km and 18.5 km altitude and  $\Delta h =$   
194 500 m. In order to minimize possible melting hail effects, we truncate horizontal reflectivity in  
195 57 dBZ such that  $Z(Z > 57 \text{ dBZ}) = 57 \text{ dBZ}$ . This calculation treats all column as liquid water  
196 for simplicity. Because convective clouds often present supercooled droplets, the separation  
197 between liquid and frozen hydrometeors can be very complex. Therefore, the VIL values  
198 reported here can be understood as an overall indication of the rain cells water mass rather than  
199 precise retrievals.

200 d. System top altitude ( $H_{top}$ , in km), calculated as the maximum echo top of 20 dBZ  
201 in the cylinder in each time step. Note that while the 34 dBZ threshold is enforced  
202 on the 2 km CAPPI for tracking purposes, the other levels do not have similar  
203 restrictions and only follow the evolution of the tracked rain cell at 2 km altitude.

204 e. VIL density (DVIL), calculated as the ratio between VIL and echo top of 20 dBZ.

205 f. Total water ( $W_T$ , in t), calculated by the areal integration of VIL.

206 Within every cell cylinder, the altitude level with the highest amount of water mass was  
207 looked for in a similar way as in Heiblum et al. (2016a,b), Chen et al. (2017) and Dagan et al.  
208 (2018). Those studies reduce the analysis of whole clouds to their COG. However, here a slightly  
209 different approach was adopted in which we do not directly calculate the mass-weighted average  
210 altitude. Instead, we define a center of activity (COA) as the altitude level that most contribute to  
211 the system average VIL. This level is somewhat similar to COG, but is much more adequate for  
212 radar applications, especially for operational applications. Because there is no mass-weighing of  
213 all altitudes, its calculation is much simplified. As will be shown later, the vast majority of the  
214 COA data is contained within the lower 4.5 km CAPPIs meaning that it is primarily related to  
215 liquid droplets for the region. The pair of COA and VIL will be used here to characterize the rain  
216 cells macroscale properties.

217 Formally, COA is calculated from (3) without the altitude integration. In this way, we  
218 have each CAPPI's contribution to the total VIL:

$$219 \quad VIL_i = 3,44 \cdot 10^{-6} (Z_i)^{4/7} \quad (3)$$

220 By averaging  $VIL_i$  for every CAPPI (i.e. every  $i$ ), we define the average vertical profile of  
221 the contributions to VIL. COA is obtained as the altitude of the level with the highest  
222 contribution.

223 For the microphysical approach, in this study we will focus on polarimetric variables and  
224 Gamma DSD parameters estimated from them. Because the DSD retrieval is an under  
225 constrained problem for radar applications, it is unfeasible to construct a full Gamma phase  
226 space as in Cecchini et al. (2017). Instead, the intent is to study DSD parameters values in the  
227 macroscale context provided by COA and VIL. We obtain the DSDs directly from the volume  
228 scans instead of the CAPPIs because  $Z_{dr}$  and  $K_{dp}$  values are also needed.

229 The Gamma DSD is estimated from the volumetric fields of  $Z$ ,  $Z_{dr}$  and  $K_{dp}$  following  
 230 Kalogiros et al. (2013). The authors show a method to obtain the normalized gamma DSD, which  
 231 has the form:

$$232 \quad N(D) = N_w f(\mu) \left(\frac{D}{D_0}\right)^\mu \exp\left[-(\mu + 3.67) \frac{D}{D_0}\right] \quad (4)$$

233 where  $N_w$  is the intercept,  $D_0$  is the mean volumetric diameter and  $\mu$  is the shape  
 234 parameter. The function  $f(\mu)$  is given by:

$$235 \quad f(\mu) = \frac{6}{3.67^4} \frac{(3.67 + \mu)^{\mu+4}}{\Gamma(\mu+4)} \quad (5)$$

236 The advantage of the Kalogiros et al. (2013) method is that it was specifically designed to  
 237 work with X-band radars, where Mie effects are significant. However, because there are three  
 238 parameters to estimate with only two DSD moments detected by the radar, the problem is under  
 239 constrained and an assumption has to be made about the relation between droplets size and DSD  
 240 shape. To constrain the equations, Kalogiros et al. (2013) propose the following  $\mu$ - $D_0$  relation:

$$241 \quad \mu = 165e^{-2.56D_0} - 1 \quad (6)$$

242 For more details on the DSD retrieval methodology see Kalogiros et al. (2013). In this  
 243 study we will use the normalized gamma format because it more suitable to radar applications.  
 244 However, it is important to point out that the normalized gamma in Equation 4 can be easily  
 245 converted to the Gamma format introduced by Ulbrich (1983) that is used in Cecchini et al.  
 246 (2017) by:

$$247 \quad N_0 = N_w f(\mu) D_0^{-\mu} \quad (7)$$

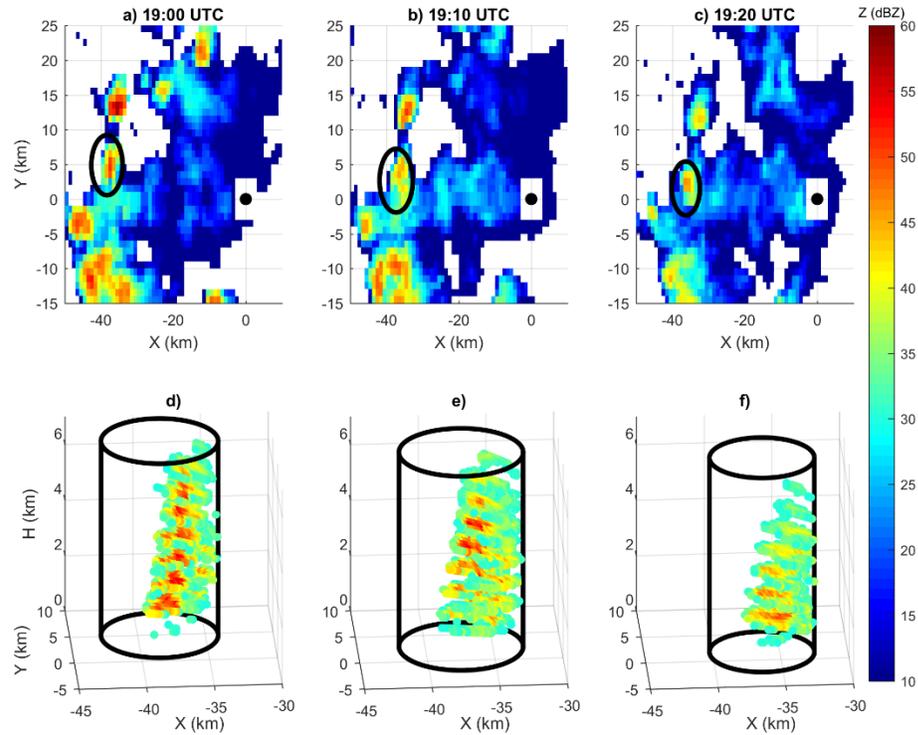
$$248 \quad \Lambda = \frac{\mu + 3.67}{D_0} \quad (8)$$

### 249 2.3 Analysis period and rain cells selection

250 The results presented here are focused on the 2016/2017 summer, more specifically  
 251 between November 22<sup>nd</sup>, 2016 and March 3<sup>rd</sup>, 2017. Within this period, 21 days were chosen for  
 252 analysis, based on data availability and occurrence of precipitation. While ForTraCC detects rain  
 253 cells throughout all radar range (100 km), we will focus exclusively on those that had their entire  
 254 lifecycle within 10 km to 60 km from the radar. Additionally, cells were manually and  
 255 qualitatively filtered based on PPI coverage and overall precipitation type. Cases associated with  
 256 small numbers of PPIs, either relatively shallow clouds or cells close to the 60 km circle that  
 257 have only a few cross-sections in the volume scan (e.g. 3 or less), or cases immersed in  
 258 stratiform systems (qualitatively analyzed case by case), were excluded. With those filters, 446  
 259 cells remained for analysis.

260 An example of the tracking algorithm functioning is shown in Figure 1. The rain cell was  
 261 a relatively shallow precipitating cloud that was detected in November 26<sup>th</sup>, 2016 at 19:00 UTC  
 262 (16:00 local time) and lasted approximately 20 minutes. The cell evolution is shown in the 2km  
 263 CAPPIs in Figures 1a-c, while the respective volume scans inside the cylinder discussed in the  
 264 previous section are shown in Figures 1d-f (limited to  $Z \geq 30$  dBZ for clarity). Note that this  
 265 strategy allows not only the study of the evolution of macroscopic characteristics such as the cell  
 266 area and depth, but also its internal structure in a Lagrangean way. Applying the same

267 methodology to all rain cells detected by ForTraCC allowed the creation of the object-oriented  
 268 dataset that will be discussed here.



269

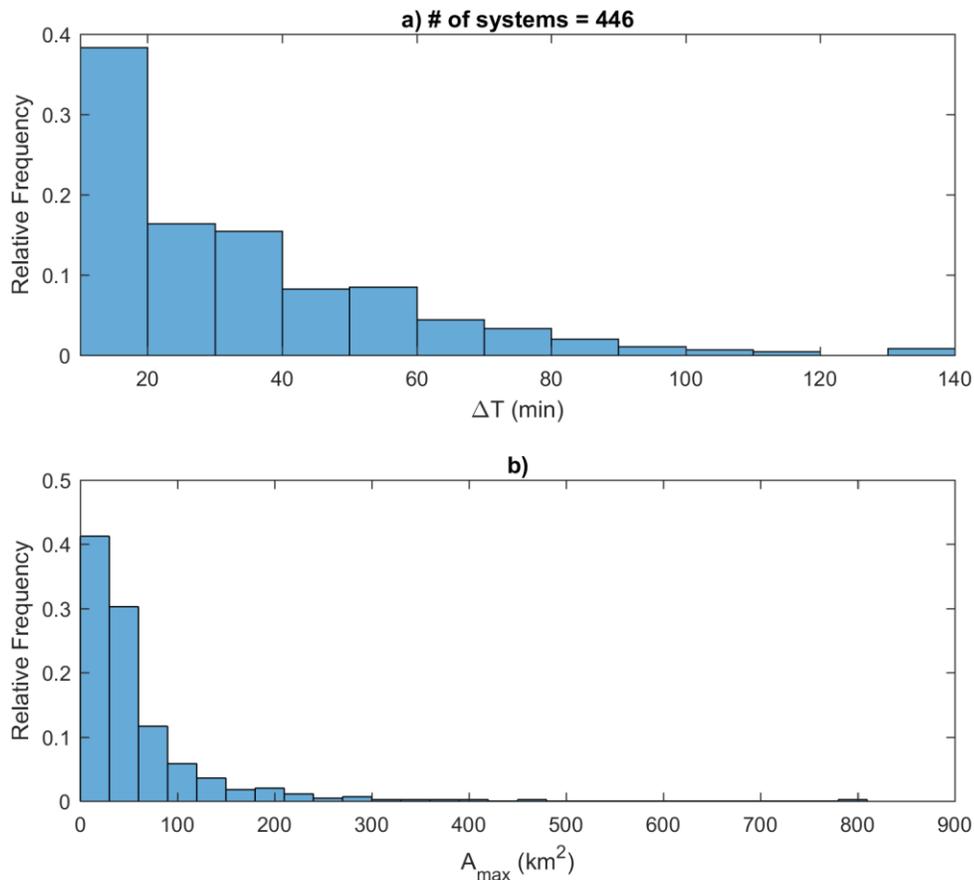
270 **Figure 1:** example of the tracking algorithm functioning. This rain cell was detected on  
 271 November 26th, 2016 at 19:00 UTC (16:00 local time) and lasted approximately 20 minutes.  
 272 Graphs in a-c are the 2-km-height CAPPI evolution of the cell (highlighted by the black circles),  
 273 as detected by the methodology described in this section. The black dot at {0, 0} is the radar.  
 274 Graphs in d-f are the respective evolution of the PPIs in the volume scan, where only  $Z \geq 30$  dBZ  
 275 points are plotted. In all graphs X and Y represent east and west distances from the radar,  
 276 respectively, and H is the altitude.

277

## 278 3 Results

### 279 3.1 Overall rain cells characteristics

280 The rain cells detected by the X-band radar throughout the analysis period present high  
 281 variability in terms of morphological aspects. While around 55% presented duration and area of  
 282 20 minutes and  $40 \text{ km}^2$  or less, respectively, there were also cells that lasted for more than one  
 283 hour that covered a few hundred  $\text{km}^2$  (Figure 2). Rain cells maximum area and duration are  
 284 usually correlated, as found by previous studies in several regions (e.g. Machado & Laurent 2004  
 285 for the Amazon). The same was found here (not shown), where the larger rain cells took longer  
 286 to dissipate.



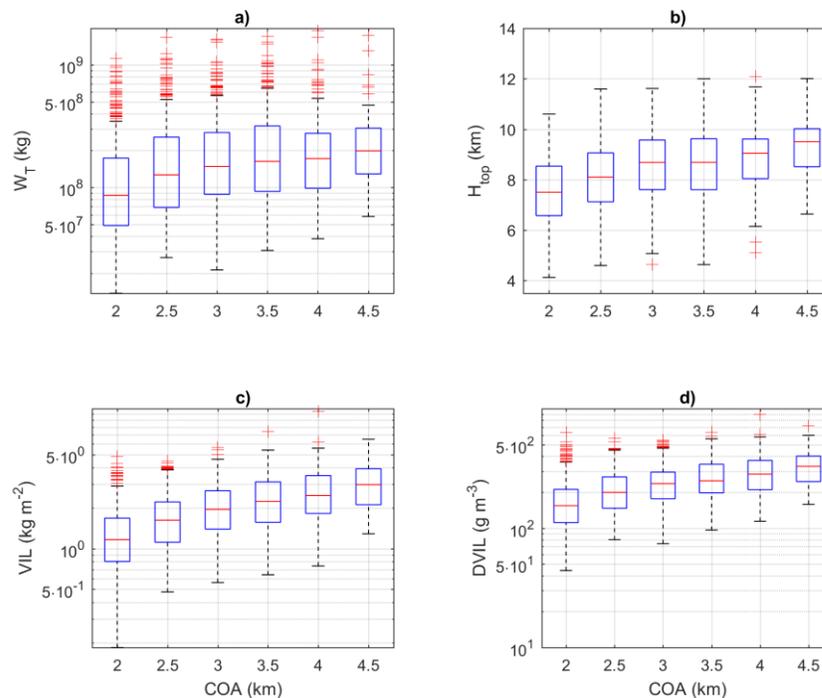
287

288 **Figure 2:** normalized histograms of a) duration  $\Delta T$  and b) maximum area  $A_{\max}$  by the cells  
 289 detected by ForTraCC. Note that the minimum values in the vertical bars are inclusive, so a bar  
 290 between 10 min and 20 min is associated to the former and so on. The same is valid for the  
 291 maximum area histogram in b).

292 In terms of the diurnal cycle of rain cells, we found two main peaks of activity: one at the  
 293 afternoon/late afternoon and another at late night/early morning. It is likely that a significant  
 294 amount of rain cells in the afternoon are linked to local convection, while events in the  
 295 night/early morning are possibly associated to larger-scale systems. Even though we do not apply  
 296 a classification criterion, it should be noted that the overall statistics shown here are largely  
 297 representative of afternoon short-lived systems because of their larger sample size (70% of the  
 298 macroscale data is within the 16:00 UTC to 22:00 UTC period, which is between 13:00 and  
 299 19:00 local time).

300 One of the main interests of this study is to analyze the property COA and how it relates  
 301 to the other characteristics of the rain cells. Figure 3 shows the relation of COA to  $W_T$ ,  $H_{top}$ , VIL  
 302 and DVIL. Firstly, we note the limitation of 500-m-resolution in the COA calculation, because of  
 303 the vertical interval between the CAPPIs. Additionally, 97% of the data present COA equal to or  
 304 lower than 4.5 km and that is why apply this limit to the figure. It also means that COA is usually  
 305 below the 0 °C level (around 4.5 km or higher for the region). This is a result not only of the  
 306 water mass accumulation on those levels but also of the different ways liquid droplets and ice

307 particles interact with radiation. Because of the oblate shape and higher dielectric constant, the  
 308 horizontal reflectivity of raindrops is higher than that of ice particles with similar sizes.  
 309 Additionally, the melting of ice results in higher reflectivity close to the 0 °C level and below.  
 310 Those factors further favor the limitation of COA to heights up to 4.5. With that said, it is  
 311 possible to note a positive relation between COA and  $W_T$  (Figure 3a),  $H_{top}$  (Figure 3b), VIL  
 312 (Figure 3c) and DVIL (Figure 3d) in terms of the median behavior. It shows that higher COA is  
 313 generally related to deeper cells with high amounts of water inside. Even though COA is related  
 314 to the median  $W_T$ ,  $H_{top}$ , VIL and DVIL, Figure 3 shows that there is a lot of variability as well. It  
 315 is possible to see the significant overlap between the interquartile ranges of different COA values  
 316 (blue bars in Figure 3). This is a result of varying rain cells characteristics such as their sizes, the  
 317 underlying thermodynamic environment they are embedded in and different life cycles that may  
 318 or may not include mergers and splits. Additionally, it highlights that COA is a characteristic that  
 319 is shared among significantly different rain cells. For instance, the interquartile range shown in  
 320 Figure 3c indicates that rain cells with COA = 4.5 km have VIL between 2 kg m<sup>-2</sup> and 4 kg m<sup>-2</sup>.  
 321 While those types of cells present significantly different quantities of water and different  
 322 intensities overall, they likely share similar vertical structure. Therefore, by combining a measure  
 323 of the cells overall amount of water with COA may prove invaluable to holistically understand  
 324 the rain cells appearance and life cycle.



325

326 **Figure 3:** boxplots of a)  $W_T$ , b)  $H_{top}$ , c) VIL and d) DVIL as function of COA.

327 The simulations shown in Heiblum et al. (2016a) show that the rain center of gravity is at  
 328 the same altitude as the cloud center of gravity when a new convective cloud is formed. When  
 329 the cloud grows and start producing more rain droplets, their center of gravity gets higher than  
 330 that of cloud droplets. The situation is reversed by the onset of precipitation. Given that all rain

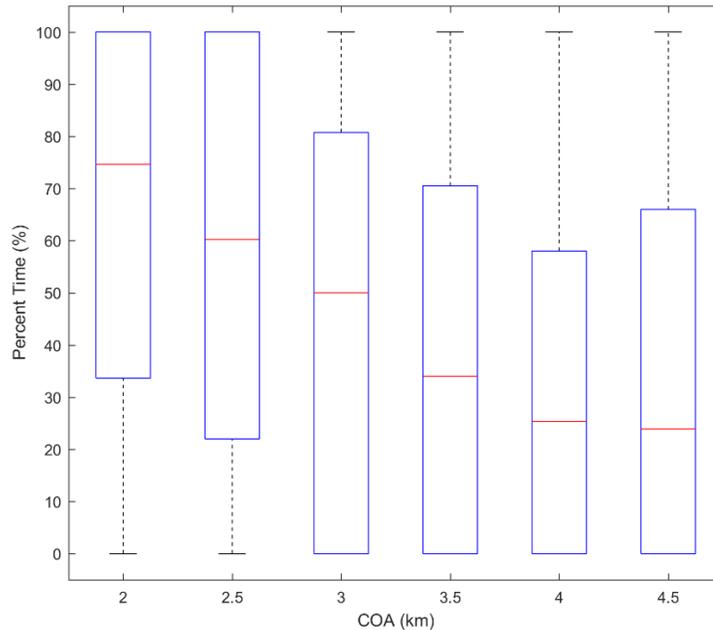
331 cells must start as shallow cumulus, a similar pattern is expected for COA. However, as  
332 discussed earlier, our methodology detects rain cells already at a relatively mature stage.  
333 Therefore, it is reasonable to expect that COA is high at the cell detection and then it sinks  
334 following the precipitation and dissipation processes. To confirm that, we calculated the rain  
335 cells percent time such that 0% corresponds to the detection and 100% to the dissipation. By  
336 averaging this time according to the cells COA, we obtained the results shown in Figure 4. We  
337 note that the 0% and 100% limits here are limited by the ForTraCC threshold chosen and are not  
338 related to the visual aspect of the clouds. Because the threshold is relatively high (34 dBZ), the  
339 results within 0% and 100% percent time cover basically the early to late mature phase. The  
340 growing and dissipation stages are likely underrepresented in our results.

341 If the ForTraCC threshold was reduced, it would likely be possible the detection of rain  
342 cells in earlier and later stages of their life cycle. However, as discussed in Section 2.2, this  
343 would also add noise to the statistics by including several smaller systems that have relatively  
344 limited precipitation capabilities and that present chaotic merger and splits events. On the other  
345 hand, high thresholds such as the one used here provide better statistics for the fewer systems  
346 detected. In this study, we chose to stick to the high threshold in order to calculate rain cell  
347 statistics without requiring too much data filtering. The downside is that the systems are detected  
348 already close to maturation and there is not much information on the growing or dissipation  
349 stage. Nevertheless, the characterization of the early and late mature stages shown here may help  
350 nowcasting techniques by providing the benchmark of where the relatively intense systems grow  
351 to. For future case studies, or even operational implementation, a lower threshold may be  
352 considered if additional filters are put in place.

353 Figure 4 confirms that the median behavior of the rain cells is to present COA around 4  
354 km or 4.5 km early on, which decays with further cell evolution. Of course, there is a lot of  
355 variability as well due to the complexities of the rain cells dynamics and the limitations of the  
356 measurement setup. Firstly, we note that different rain cells may present different COA motions  
357 according to their thermodynamics. For instance, cells that are able to sustain strong updrafts  
358 may support higher COA for longer periods of time. On the other hand, cells with high  
359 precipitation efficiency can have rapidly sinking COA. Additionally, the 10 minutes limitation of  
360 the CAPPIs means that the cells may be captured at slightly different stages of their relative life  
361 cycle.

362 In order to further understand the role of COA, it is important to analyze the cell  
363 characteristics discriminating by its value. Figure 5 presents reflectivity contoured frequency by  
364 altitude diagram (CFAD, Yuter & Houze, 1995) plot for  $Z > 10$  dBZ when the cells present COA  
365 between 2 km and 4.5 km. Here we note that, while ForTraCC was set up to work with a 34 dBZ  
366 threshold, lower reflectivity values are contained in the rain cells cylinders (as in Figure 1) for  
367 two reasons. Firstly, the 34 dBZ limit is only enforced for the 2 km CAPPI, so lower  $Z$  can  
368 appear in higher levels. Additionally, the added 2 km in the rain cells radii described in Equation  
369 1 means that lower  $Z$  values are allowed inside rain cells even at the 2 km altitude level. The  
370 CFADs in Figure 5 were obtained from the CAPPIs, with  $H$  and  $Z$  intervals of 0.5 km and 2.5  
371 dBZ, respectively. The number of data in each bin is represented in colors, while the dashed  
372 black lines are averaged  $Z$  profiles. The horizontal lines represent the respective COA value for  
373 clarity. This figure clearly shows that COA has a strong relation to the overall appearance of the  
374 rain cells. Note that all graphs in Figure 5 show a boomerang-like shape (except Figure 5f where  
375 it is cutoff) with COA at the middle. It shows that the cells usually have the appearance of

376 Figures 5a,b when they are detected and then evolve along the sinking COA (Figures 5c-f),  
 377 taking into account the results shown in Figure 4.

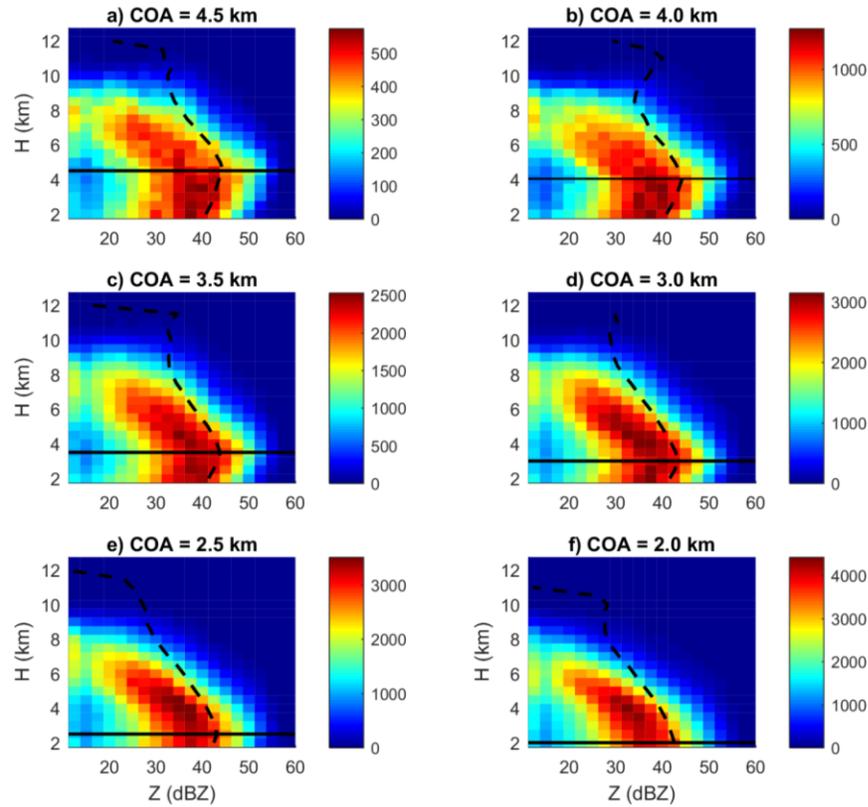


378

379 **Figure 4:** boxplot of percent time as function of COA. Percent time is 0% on system formation  
 380 and 100% on dissipation according to the ForTraCC criteria chosen.

381 The boomerang-like shape can be mostly explained by a few competing processes. For  
 382 high COA (e.g.  $\geq 3.5$  km), the decrease of  $Z$  with altitude is characteristic of growing ice-to-  
 383 liquid ratios. The higher the relative amount of ice, the lower the reflectivity will be. For lower  
 384 COA, the decrease in  $Z$  is related not only to the ice-to-liquid ratio but also to the overall rain cell  
 385 collapse. Below COA, the  $Z$  decrease towards the surface is most likely associated to  
 386 evaporation. However, collection processes also take place in regions close and below the COA.  
 387 Because COA is associated to the highest reflectivity values, it means that collection processes,  
 388 together with ice melting, dominate in the COA layer.

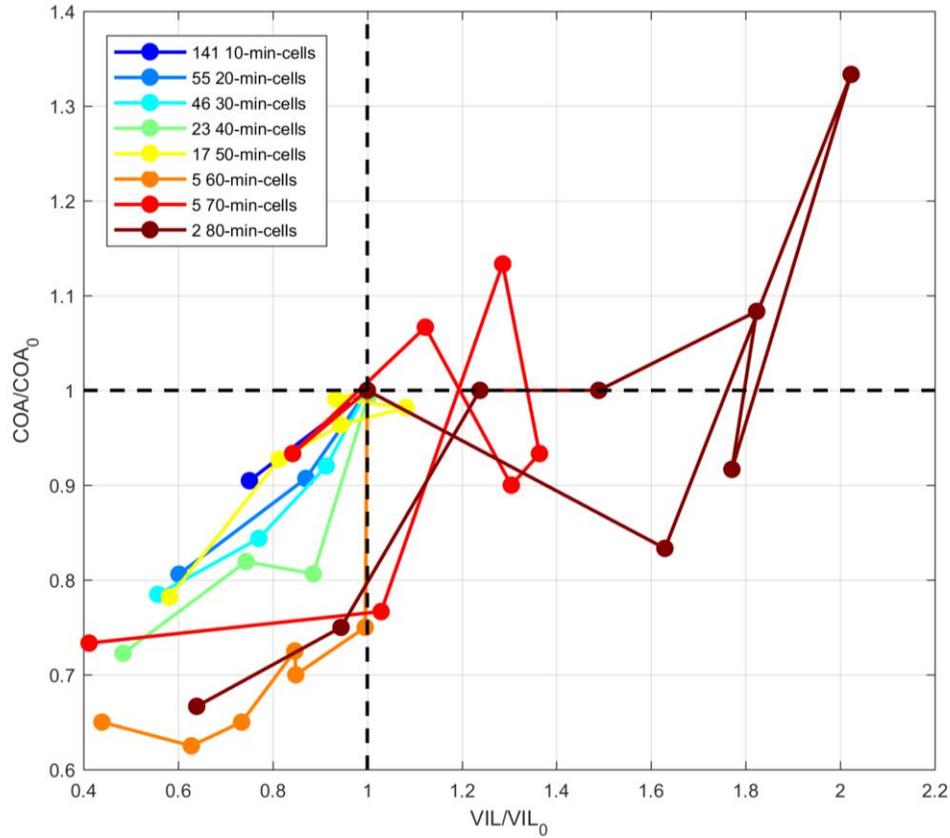
389 As mentioned earlier, there is significant variability of reflectivity values and  
 390 consequently water mass for systems sharing the same COA. Therefore, despite having different  
 391 absolute values, it is possible to conclude that COA captures one fundamental aspect of the rain  
 392 cells. But it is noteworthy that the variability of  $Z$  around COA is similar for all values of the  
 393 latter. Around COA the reflectivity values tend to vary between 30 dBZ to 45 dBZ in terms of  
 394 the most common rain cells observed here. Additionally, the averaged  $Z$  at COA varies only  
 395 roughly 2 dBZ, going from 44.5 dBZ in Figure 5a down to 42.6 dBZ in Figure 5f and decaying  
 396 almost linearly with COA. This indicates that COA is a layer in relative balance not only in the  
 397 vertical but also throughout the rain cells life cycles. Below COA, the averaged profiles are  
 398 within a range of 3.7 dBZ and this difference tends to grow with altitude. For instance, above 8  
 399 km the maximum difference between the averaged profiles reach 29.4 dBZ due to high  
 400 reflectivity for  $3.5 \text{ km} \leq \text{COA} \leq 4.5 \text{ km}$  and low reflectivity for  $2.0 \text{ km} \leq \text{COA} \leq 3.0 \text{ km}$ . This  
 401 shows that higher COA are associated to more active ice phase given the likely higher water  
 402 mass available at high altitudes. It is also likely that high COA are related to  $Z_{dr}$  columns.



403

404 **Figure 5:** reflectivity CFADs discriminated by COA. The continuous black lines represent COA  
 405 for clarity, while the dashed black lines represent averaged Z profiles for the same data  
 406 (calculated on linear scale before converting back to dBZ). The number of points in each H and  
 407 Z bins are shown in colors.

408 While Figure 5 paints a cohesive view of the rain cells, it lacks representation of  
 409 individual systems that may have different COA physics. For instance, short- and long-lived  
 410 cells can present contrasting COA characteristics because the dynamic and thermodynamic  
 411 processes sustaining them is different. In order to address this issue, we present the results in a  
 412 similar way to Heiblum et al. (2016a) in Figure 6. The intent of this calculation is to obtain the  
 413 average behavior of cells sharing similar durations. Because each system may have different  
 414 characteristics at detection, a normalization process is used. The evolutions shown in Figure 6  
 415 are all relative to the properties observed at detection (identified with a 0 subscript). Because of  
 416 the normalization, all trajectories start at the (1,1) point, indicated by the junction of the two  
 417 dashed black lines. From this point on, the trajectories are calculated for the same 10-minute  
 418 time step of the cells sharing the same duration – 10-minutes cells have two points, 20-minutes  
 419 cells present three points and so on. Every trajectory represents averaged values between cells of  
 420 the same duration, whereas cells with mergers and/or splits detected by ForTraCC were  
 421 excluded.



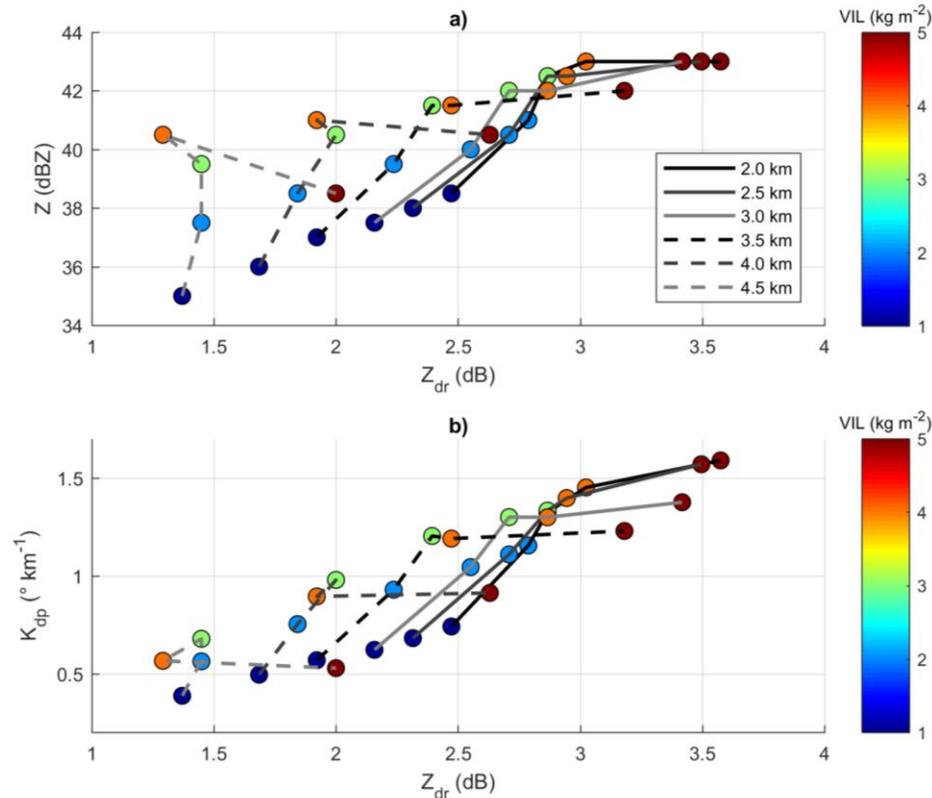
422

423 **Figure 6:** trajectories in the normalized phase-space COA-VIL. Both COA and VIL are  
 424 normalized by their values at system detection ( $COA_0$  and  $VIL_0$ , indicated by black dashed lines)  
 425 to compensate for possible different initial conditions. All trajectories start at the (1,1) point and  
 426 evolve from there. The number and duration of the systems is given in the legend.

427 Figure 6 shows that cells lasting up to 60 minutes have a similar pattern in the COA-VIL  
 428 phase space. From the starting point, they evolve towards the third quadrant, where both COA  
 429 and VIL diminish over time. The difference between their trajectories is mainly on how they  
 430 evolve in this quadrant. Except for the yellow line (50-minutes cells), there is a trend of  
 431 increasing curvature from short- to long-lived cells. In other words, short-lived cells tend to lose  
 432 COA and VIL in a similar proportion, whereas longer-lived ones may retain some VIL even  
 433 while COA sinks. This can either mean that longer-lived cells usually start with higher COA or  
 434 that the layer around their COA contain relatively more water that remains in the cloud when the  
 435 bottom part precipitates. Even though there is a lack of a significant sample size for cells lasting  
 436 70 minutes or more (with no merger and/or split), it is notable that they were the only ones that  
 437 meaningfully evolved through the right quadrants. Those quadrants exemplify processes in  
 438 which the rain cells gain mass with either ascending or descending COA. Therefore, longer-lived  
 439 cells present some sustained source of water mass even after it reached the relatively mature  
 440 stage as detected by our methodology. This could be either microphysical mechanisms, such as  
 441 continued droplet growth by sustained updrafts or melting ice falling from above, or some form  
 442 of dynamical feedback. Either way, a case-by-case analysis would be ideal to elucidate such  
 443 aspects.

444 Given that the COA is a layer of interest because of its relation to rain cell life cycle and  
 445 internal structure, it is important to analyze its microphysical properties to provide a reference  
 446 for the comparison of individual cases. Figures 7 and 8 show median polarimetric variables and  
 447 associated DSD parameters taking into account the macro characteristics of the cells (COA and  
 448 VIL). Every point represents median properties from volume scans in 1-km vertical layers  
 449 around COA and for every  $1 \text{ kg m}^{-2}$  VIL interval (see figure description for more details), for the  
 450 same cells as in Figure 6. Basically, it represents a snapshot of the most intense region within the  
 451 COA level.

452



453

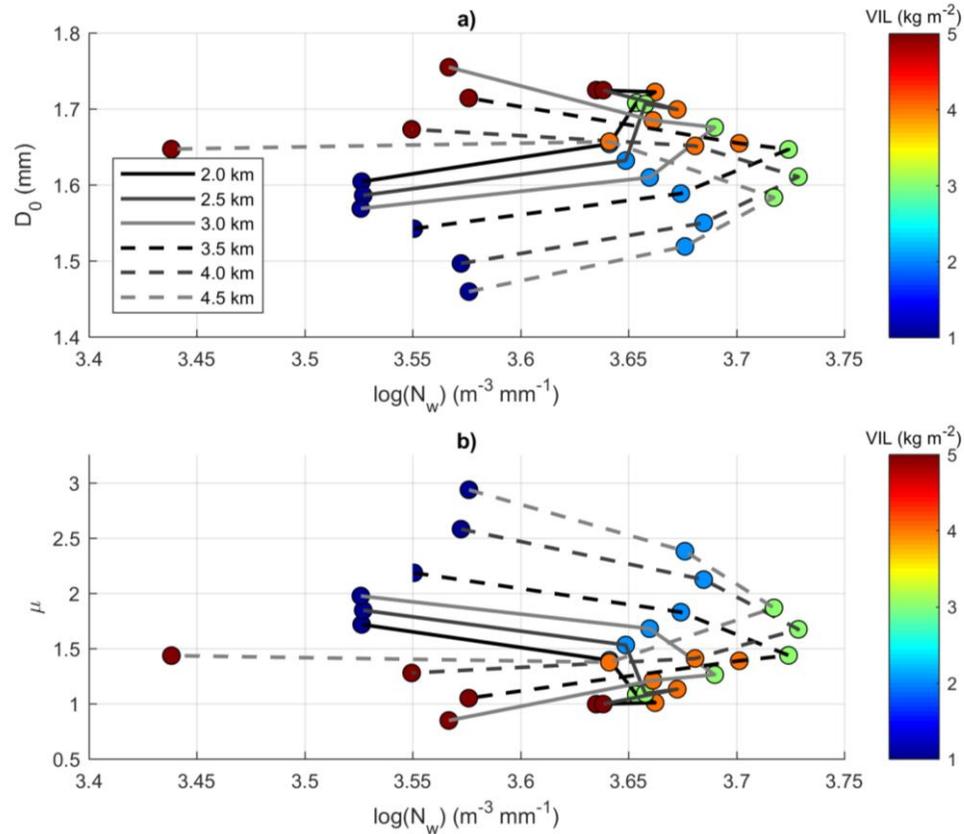
454 **Figure 7:** median polarimetric variables as function of COA and VIL. While VIL and COA were  
 455 obtained from the CAPPIS, Z, Z<sub>dr</sub> and K<sub>dp</sub> are from the volume scans, limited by the following  
 456 thresholds: 1)  $Z > 30 \text{ dBZ}$ ; 2)  $Z_{dr} > 0.5 \text{ dB}$ ; 3)  $K_{dp} > 0 \text{ ° km}^{-1}$ ; 4)  $\rho_{HV} > 0.97$ ; 5) elevation  
 457 angle  $\alpha \leq 15 \text{ °}$  and 6)  $\text{COA} - 0.5 \text{ km} \leq H \leq \text{COA} + 0.5 \text{ km}$ . Those thresholds aim to focus the  
 458 analysis primarily on the cell core at the 1 km layer around COA. The rain cells considered here  
 459 are the same as in Figure 6, where no merger or split was automatically detected. VIL values are  
 460 represented by circle colors, showing the upper bound of every  $1 \text{ kg m}^{-2}$  interval. For instance,  
 461 dark blue points are shown as  $\text{VIL} = 1 \text{ kg m}^{-2}$  and are related to the interval  $0 \text{ kg m}^{-1} < \text{VIL} \leq 1$   
 462  $\text{kg m}^{-2}$ . The different curves represent medians of all volume scan measurements that follow the  
 463 criteria above while the cell presented the specified COA (legend in Figure 7a). The Z and Z<sub>dr</sub>  
 464 medians were obtained directly in dB scale. The average number of data for each point is 23366,  
 465 going from a minimum of 2387 to a maximum of 71837.

466

467 Figure 7 shows the evolution of  $Z$ ,  $Z_{dr}$  and  $K_{dp}$  according to COA and VIL individually,  
 468 as well as a combination between them. While an increase of the polarimetric variables with VIL  
 469 is expected, it is somewhat counter-intuitive that they also increase with decreasing COA  
 470 because of rain cell dissipation. As they approach 2 km in altitude, the cells are closer to being  
 471 dissipated but their inner core have reached peak values of the polarimetric variables. However,  
 472 it is important to highlight that this will only happen in case VIL remains relatively constant. In  
 473 individual rain cells, both COA and VIL can diminish during dissipation (Figure 6) and the  
 474 overall polarimetric characteristics will largely depend upon their balance.

475 Another interesting pattern in Figure 7 is the capping in  $Z$  and  $K_{dp}$  values for high VIL.  
 476 For fixed COA levels,  $Z$  and  $K_{dp}$  tend to grow proportionally to  $Z_{dr}$  up to  $VIL = 3 \text{ kg m}^{-2}$ . From  
 477 this point on,  $Z_{dr}$  continues to grow while both  $Z$  and  $K_{dp}$  either stagnate or increase ever so  
 478 slightly ( $Z$  even decreases from  $VIL = 4 \text{ kg m}^{-2}$  to  $VIL = 5 \text{ kg m}^{-2}$  for  $COA = 4.5 \text{ km}$  and  $COA =$   
 479  $4.0 \text{ km}$ ). This indicates that high VIL cells present different microphysical characteristics in its  
 480 center of gravity. Those cells have both bigger droplets overall (Figure 8) and larger amounts of  
 481 water for the sinking-COA-droplets to collect, further favoring their growth. However, possible  
 482 attenuation of the radar signals can also help explain the capping – especially in  $Z$ .

483 From the results shown in Figure 7, the corresponding normalized gamma DSD parameters were  
 484 obtained by applying the method described in Kalogiros et al. (2013). This method was  
 485 developed to correct for Mie effects specifically for the X-band range and the results are shown  
 486 in Figure 8. Firstly, we note that the mirroring between Figure 8a and Figure 8b is explained by  
 487 the  $\mu$ - $D_0$  constrain in Equation 6. Nonetheless, the range of  $\mu$  values provided can be used to  
 488 further characterize the cells COA microphysical properties. Following the patterns in Figure 7,  
 489  $D_0$  also tends to increase with increasing VIL and decreasing COA. However, this is not  
 490 followed by a monotonic variation in  $N_w$ . Note that this parameter tends to grow from  $VIL = 1 \text{ kg}$   
 491  $\text{m}^{-2}$  up to  $VIL = 3 \text{ kg m}^{-2}$  and then decreases with increasing VIL. When both  $N_w$  and  $D_0$   
 492 increase, the overall number of droplets of the DSD increases unevenly in terms of contribution  
 493 to the total water mass. In other words,  $D_0$  increases by an increase of the number of relatively  
 494 big droplets even when smaller droplets are also more numerous. On the other hand, lower  $N_w$   
 495 with bigger  $D_0$  indicate that there are less droplets on both ends, but the smaller ones contribute  
 496 even less to the DSD water mass. That difference may be explained by the increased efficiency  
 497 of the collection process on high VIL cases, where a turning point is around  $VIL = 3 \text{ kg m}^{-2}$ . This  
 498 turning point approximately coincides with the  $Z$  and  $K_{dp}$  capping shown in Figure 7.



499

500 **Figure 8:** same as Figure 7 but for the DSD parameters obtained by the respective medians of Z,  
 501 Zdr and Kdp using the method described in Kalogiros et al. (2013).  $N_w$  is shown in log-scale  
 502 (base 10) for simplicity.

503 3.2 Case studies

504 Two case studies were selected to exemplify physical characteristics of individual rain  
 505 cells as seen by the COA/VIL approach. The cells were selected not based on their particular  
 506 intensity, but rather to exemplify how different COA characteristics are related to the overall  
 507 cells appearance. We chose two rain cells during November 28<sup>th</sup>, 2016, with overall properties  
 508 shown in Table 1. For the case studies selection, we allowed merger and splits occasions because  
 509 it was possible to analyze them in more detail.

510

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514

515 **Table 1:** overall characteristics of the case studies. \*Calculated as the sum of accumulated  
 516 rainfall in every 2-km-CAPPI pixel with  $Z > 10$  dBZ using the Marshall-Palmer Z-R relation and  
 517 assuming that each pixel remained static ( $Z = \text{constant}$ ) for the 10-minute intervals.

Case # and time steps (UTC)	COA (km)	VIL ( $\text{km m}^{-2}$ )	$H_{top}$ (km)	New/Split/Merger/ Continuity	Total Accumulated Rainfall* (mm)	A ( $\text{km}^2$ )
<b>Case 1</b>						
17:40	4.5	6.49	9.0	New	78	14
17:50	3.5	2.49	8.6	Continuity	267	41
18:00	4.0	2.08	9.0	Continuity	404	36
18:10	3.5	3.41	9.5	Merger	763	133
18:20	3.5	2.60	9.1	Continuity	1014	85
18:30	3.5	1.76	9.1	Split	1131	34
<b>Case 2</b>						
18:50	4.5	2.41	6.6	New	73	18
19:00	2.5	2.13	7.0	Continuity	209	22
19:10	2.0	1.04	7.0	Continuity	302	26
19:20	2.0	0.74	7.0	Merger	350	19
19:30	2.0	0.86	7.0	Continuity	413	20
19:40	2.0	0.66	7.1	Continuity	480	27

518

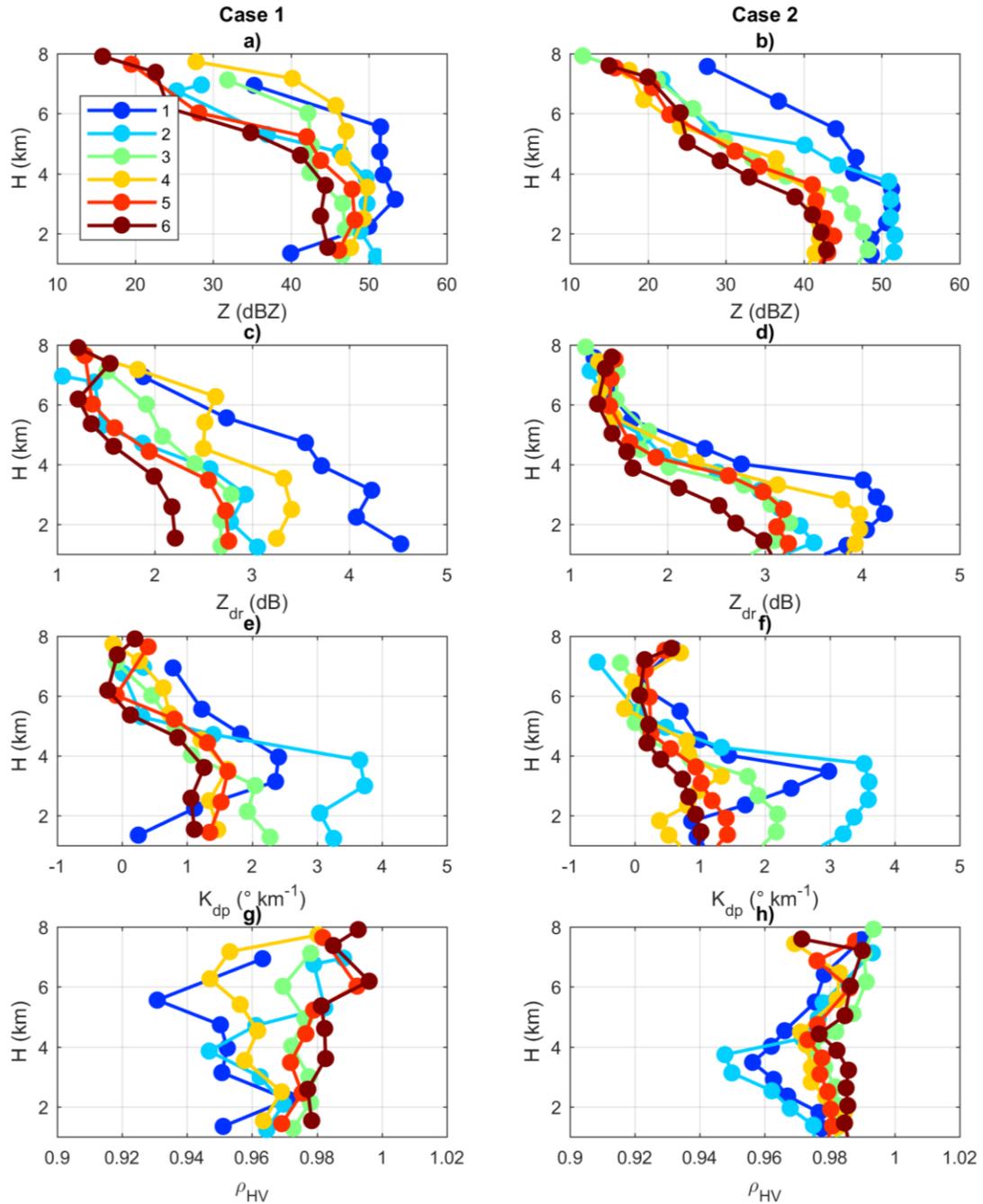
519 From Table 1, we note that the rain cells followed the overall trend of high COA upon  
 520 detection. However, we chose two cases with different COA trends. Note that Case 1 presents  
 521 relatively high COA throughout its lifecycle, which ends in a split (the resulting cells from the  
 522 split did not meet the ForTraCC requirements to be tracked). On the other hand, Case 2 present  
 523 a rapidly decaying COA. Both cases presented merger occasions automatically detected by  
 524 ForTraCC. The merger in Case 1 was followed by an increase in VIL, A and  $H_{top}$ , but the same  
 525 was not observed for Case 2. This could be a reflection of the different convective environment  
 526 for the two cases, illustrating the complexities of the dynamics involved. From Table 1 it is clear  
 527 that Case 1 is the stronger one, with more than double the total accumulated rainfall of Case 2.

528 For each case study selected, the cells cores were looked for based on polarimetric  
 529 variables thresholds. For every radar elevation angle, we have a respective cross-section (slice)  
 530 of the rain cells. In each of those slices, the core is found by: 1)  $Z \geq Z_{0.75}$  for altitudes below 4.5  
 531 km, where  $Z_{0.75}$  is the 75% percentile of  $Z$  in the slice and 2)  $Z \geq 10$  dBZ,  $Z_{dr} \geq 1.0$  dB and  $\rho_{HV} \geq$   
 532 0.85 for altitudes above 4.5 km and below 8 km (based on the  $Z_{dr}$  column criteria of Carlin et al.,  
 533 2017). The 4.5 km altitude is an approximation for the 0 °C isotherm level. Given that we are

534 analyzing only a few cases in more detail and that criteria 1 and 2 perform similarly for liquid-  
535 dominated volumes, there was no need for precise measurements of 0 °C isotherm altitude and  
536 the value is close to average radiosonde measurements. For the two criteria, we further enforced  
537 horizontal consistency as in Carlin et al. (2017), where every 9x9 range-azimuth pixel box was  
538 required to have at least 5 pixels meeting requirements 1 or 2. Overall, the criteria aim to find the  
539 most active portion of the systems along radar rays and where liquid droplets predominate.

540 Figure 9 shows the evolution of vertical profiles of polarimetric variables for the case  
541 studies, where each point represents averaged properties based on the criteria mentioned above.  
542 We chose to use averages instead of medians (as in Figure 7) because there were no rigid spatial  
543 criteria, i.e. we do not calculate averaged properties inside a specific cartesian volume.  
544 Additionally, because we are now treating each cell individually, the overall distances from the  
545 radar are similar and so are the illumination volumes of each pixel considered (which is not the  
546 case in Figure 7, where different cells are considered). In that way, all pixels are considered to  
547 have the same weight to the average in a reasonable approximation. For both  $Z$  and  $Z_{dr}$ , the  
548 averages were taken in linear scale and then converted back to dBZ and dB, respectively.

549 The rain cells usually start off with the highest  $Z$  and  $Z_{dr}$  values overall (Figure 9), which  
550 then decays as COA sediments. The rate in which  $Z$  diminishes is higher in the upper parts of the  
551 systems (e.g. above 5 km) as COA moves away from those layers. It is interesting to note that  
552 peak  $K_{dp}$  values in the lower levels occur 10 minutes after the peak in  $Z_{dr}$  at system detection.  
553 This might be explained by the different balance between droplets mean size and their overall  
554 number concentration. Peak  $Z_{dr}$  values are associated to peak  $D_0$ , but there are relatively few  
555 droplets as seen by low  $N_w$  values. This illustrates the process in which the biggest droplets reach  
556 the ground faster given their higher terminal velocity and are then followed by the bulk of the  
557 precipitation. Indeed, Table 1 shows that the total accumulated rainfall increases by a factor of  
558 2.9 (Case 2) and 3.4 (Case 1) from the first to the second time steps of the cells, which is the  
559 highest increase during their lifecycle. Of course, this increase is due not only to increased  
560 precipitation rates but also to increased covered area at 2 km altitude. Additionally, a portion of  
561 the  $K_{dp}$  peak can be explained by the melting of ice as the upper layers start to collapse, which  
562 should also decrease  $Z_{dr}$ .



563

564 **Figure 9:** averaged vertical profiles of polarimetric variables for Case 1 (left column) and Case 2  
 565 (right column). Each datapoint represents the average of a single slice of the system, according to  
 566 the criteria mentioned in the text. The legend specifies the time-steps of each case in  
 567 chronological order (same number of time steps for both cases, where the specific UTC times are  
 568 given in Table 1).

569 Averaged  $\rho_{HV}$  profiles are provided in Figure 9 as a general indication of the level of  
 570 hydrometeor mixture in the rain cells cores. We note that  $\rho_{HV}$  starts lower at cell detection  
 571 throughout the core, suggesting higher levels of mixture, and then converges to higher values

572 (around 0.98) for later stages of the lifecycle. This shows that when the cells are detected it is  
 573 more likely to find both ice particles below and liquid droplets above the 4.5 km altitude mark.  
 574 Also noteworthy is that the values are much higher than the limit of 0.85 imposed in the filters.  
 575 The overall appearance of the  $\rho_{HV}$  profile presents correlations with the systems macroscale  
 576 characteristics. Note that Case 2 have much smoother  $\rho_{HV}$  profiles as compared to Case 1 with  
 577 lower values contained in the region between 2.0 km and 4.0 km. This pattern is followed by a  
 578 characteristic variation of  $Z_{dr}$  where it quickly decays with both time and altitude above the 3 km  
 579 mark. Both mechanisms are related to the relatively quick sedimentation of COA. As shown in  
 580 Table 1, Case 1 have COA at 2.5 km already at the second time step, limiting its ability to mix  
 581 ice and liquid hydrometeors and consequently to generate a strong (or large)  $Z_{dr}$  column. On the  
 582 other hand, Case 1 maintain relatively high COA for a longer period of time which is  
 583 accompanied by higher  $Z$  and  $Z_{dr}$  (and even  $K_{dp}$ ) above 4.5 km altitude. This indicates that COA  
 584 may be a good indicator of  $Z_{dr}$  columns in continental systems, especially when paired with VIL  
 585 estimates or  $Z$  profiles. Additionally, Lier-Walqui et al. (2016) show that positive  $K_{dp}$  values  
 586 above the melting level is usually associated to updrafts meaning that this is likely one of the key  
 587 mechanisms to sustain high COA in Case 1. If high COA is considered to be a proxy for both  
 588 updrafts and  $Z_{dr}$  columns, this could be an important contribution for operational purposes given  
 589 that the identification of COA does not rely on polarimetric retrievals.

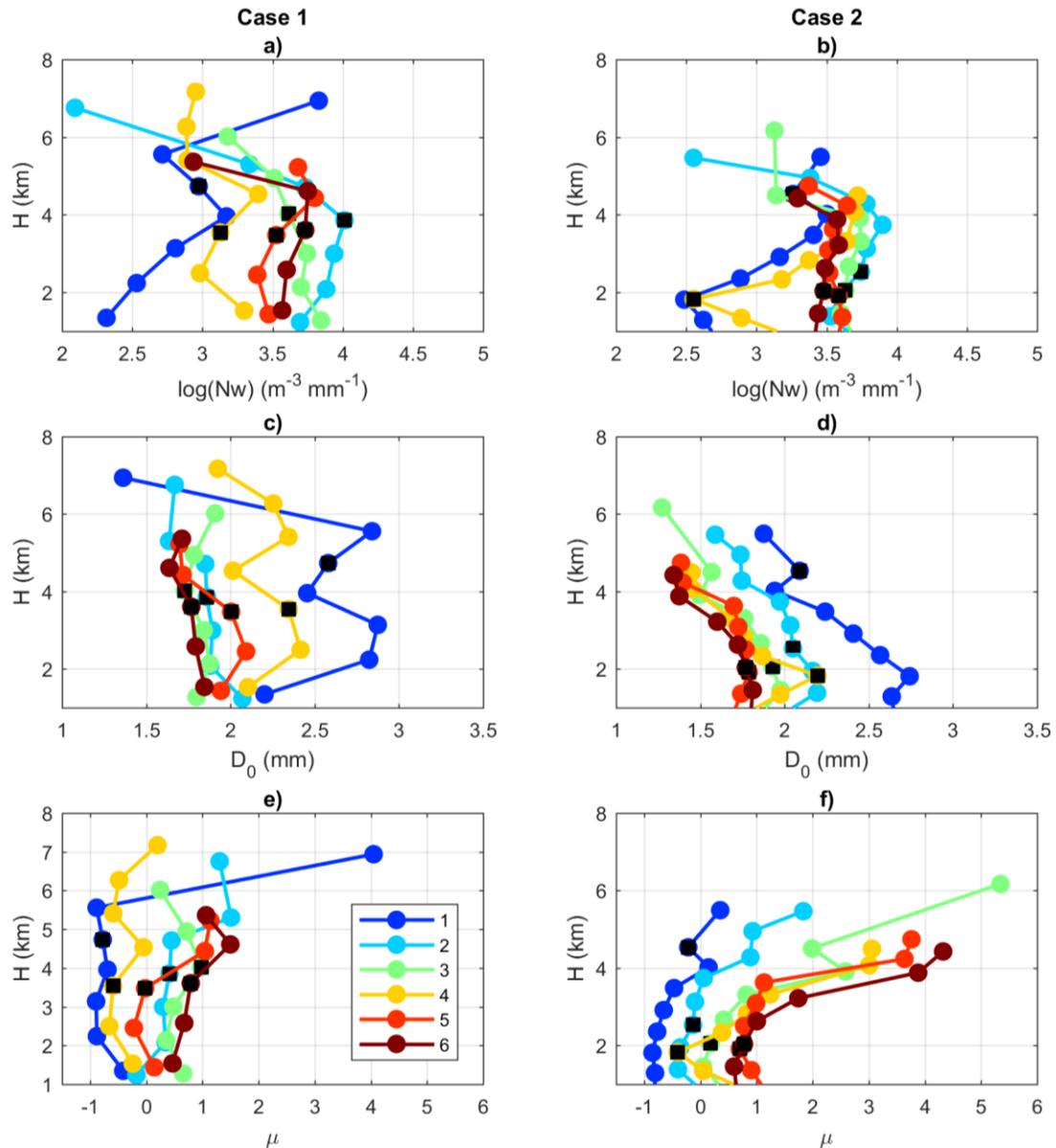
590 The polarimetric profiles for Case 1 resemble profiles of highly electrically active  
 591 systems shown in Mattos et al. (2016), where they present high  $Z$ ,  $Z_{dr}$  and  $K_{dp}$  with low  $\rho_{HV}$  in the  
 592 low levels. At higher levels, those systems have slightly positive  $K_{dp}$  and  $Z_{dr}$  accompanied by  $Z >$   
 593 30 dBZ. This kind of values were observed for Case 1 for several time steps in the beginning of  
 594 the life cycle captured by our methodology. At later stages of development,  $Z$  falls below the 30  
 595 dBZ threshold above 4.5 km, which is followed by a decrease in  $K_{dp}$ . According to Mattos et al.  
 596 (2017), positive  $K_{dp}$  values are observed slightly before the first lightning, indicating that our  
 597 case study was likely captured around that time.

598 The Gamma DSD parameters (Equation 4) can provide further insights into the  
 599 microphysical characteristics of the rain cells cores. The DSD parameters shown in Figure 10  
 600 were obtained from the data of Figure 9 with the following additional filters: 1) elevation angles  
 601 between  $1.8^\circ$  and  $15^\circ$  in order to minimize both ground noise and effects of high elevation angles  
 602 on  $Z_{dr}$  and  $K_{dp}$ ; 2)  $\mu \leq 6$ ; and 3)  $2 \leq \log(N_w) \leq 6$ . DSD parameters outside the range of filters 2  
 603 and 3 are primarily associated to higher altitudes and are not central to the discussions here.

604 As mentioned earlier, the DSD parameters indicate that droplets reach maximum size  
 605 upon the cell detection, which is accompanied by relatively small concentrations as seen by low  
 606  $N_w$ . From this point on, the overall trend is for increasing  $N_w$  and  $\mu$  with decreasing  $D_0$ . The  
 607 positioning of the COA seems to have an impact in the overall DSD parameters profiles, not  
 608 limited only to its layer. A good example can be seen in Figure 10e for Case 1. We note that  
 609 when COA is high,  $\mu$  remains relatively constant with altitude. However, when COA starts  
 610 decaying,  $\mu$  may increase for  $H > COA$ . This is enhanced in the latter stages of Case 2 because of  
 611 the faster COA sedimentation. Therefore, systems that are able to maintain high COA are likely  
 612 to have relatively big droplets throughout its core ( $\mu$  is inversely proportional to  $D_0$ ) and a strong  
 613  $Z_{dr}$  column, which is also favored when VIL is high (see Table 1 for Case 1).

614

615



616

617 **Figure 10:** Gamma DSD parameters calculated from the averaged profiles shown in Figure 9 for  
 618 Case 1 (left column) and Case 2 (right column). Black squares represent the closest datapoint to  
 619 COA for each time step. Chronological time steps are given in the legend of panel e, where the  
 620 respective UTC times are given in Table 1.

621 In terms of the DSDs close to the COA along the rain cells lifecycle (black squares in  
 622 Figure 10), we note that the case studies presented significantly lower  $\mu$  and higher  $D_0$  as  
 623 compared to the medians shown in Figure 8 (taking into account VIL and COA values). The  
 624 comparison for  $N_w$  is not as straightforward because of the change in behavior around  $VIL = 3 \text{ kg}$   
 625  $m^{-2}$ , but the case studies usually present more extreme values, be it to the lower or higher end.  
 626 This observation highlights the difference between the treatment of the systems as a bulk or  
 627 individually. The overall medians in Figure 8 are necessarily biased towards the most common  
 628 characteristics, which are often related to relatively weaker cells within the COA and VIL

629 intervals. Therefore, Figure 8 can now be understood as an approximation of the minimum  
630 properties a system must have in order to generate enough precipitation at 2 km altitude to be  
631 detected by a tracking algorithm such as ForTraCC.

632 One point in common between Cases 1 and 2 is the presence of negative  $\mu$  values at COA  
633 at detection. As discussed earlier, this point is probably close to the lightning activity initiation  
634 and the formation of  $Z_{dr}$  columns. Therefore, the evolution of  $\mu$  at COA throughout the cells  
635 lifecycle could also be considered for nowcasting applications.

636

#### 637 **4 Conclusions**

638 In this study, the macro- and microphysical characteristics of the 2016/7 summer rain  
639 cells were explored as part of the SOS-CHUVA experiment. The cells were tracked by the  
640 ForTraCC algorithm allowing for the analysis of Lagrangean properties. The cells studied were  
641 primarily afternoon convective systems with average duration and area of 29 minutes and 59  
642  $\text{km}^2$ , respectively. However, some cells lasted longer and covered larger areas – having 90%  
643 percentiles of 1 hour and 127  $\text{km}^2$ , respectively. For each cell the calculation of the vertical  
644 center of activity (COA) was introduced as a mean to locate the most active layer of the cells in  
645 terms of water mass.

646 It was shown that COA varies mostly between 2 km and 4.5 km and presents positive  
647 relations to total cell water ( $W_T$ ), VIL and DVIL. In order to further understand COA and its  
648 relation to the rain cells development, an exploratory analysis was applied to detail its relation to  
649 precipitation formation and system development and microphysics.

650 When the cells are first detected by ForTraCC (using a threshold of 5  $\text{mm h}^{-1}$  or 34 dBZ  
651 in the 2 km CAPPI), COA is usually around 4 km or 4.5 km. At this point, our results suggest  
652 that they are close to their mature stage, right before a sudden increase in the precipitation  
653 amounts at 2 km (the lower CAPPI used). Therefore, it shows that mature rain cells usually  
654 present high COA, which sediments as they collapse and dissipate. Averaged reflectivity profiles  
655 showed that  $Z$  is around 44.5 dBZ at the COA level when it is at 4.5 km, diminishing only by 2  
656 dBZ when the cells dissipate with COA at 2.0 km. Therefore, it is possible to follow the bulk of  
657 the water within the rain cells by tracking the COA evolution, which is a level of relative balance  
658 between collection and evaporation processes.

659 The COA level was found to be a property shared by rain cells with varying  
660 characteristics such as size, average VIL or echo tops. Therefore, it is interesting to add other  
661 criteria to study the rain cells. In this study we used VIL together with COA to show overall and  
662 specific characteristics of the cells measured during SOS-CHUVA. Contrary to the results  
663 mentioned above, we found that the reflectivity at the cells core and at the COA level tend to  
664 increase with decreasing COA, given that VIL remains relatively constant. The same was  
665 observed for  $Z_{dr}$  and  $K_{dp}$ , meaning that droplets may continue to grow on the cells cores even  
666 during the dissipation stage provided that there is enough background water to support their  
667 growth. On the other hand, cells sharing the same COA can present enhanced polarimetric  
668 characteristics the higher their VIL is. Therefore, the evolution of observed cells can be  
669 understood by the balance between COA and VIL variability. The relatively small variability on  
670 the overall averaged  $Z$  at COA (i.e. the 2 dBZ mentioned earlier) indicates that the balance

671 between collection and evaporation processes can also be visualized in the VIL-COA approach  
672 for the cells measured during SOS-CHUVA.

673 The microphysical characteristics of the rain cells cores were also analyzed by the  
674 respective Gamma DSD parameters. It was shown that the shape parameter  $\mu$  have an overall  
675 median variability between 0.5 and 3.0, associated to a  $1.4 \text{ mm} \leq D_0 \leq 1.8 \text{ mm}$  interval. The  
676 general trend is for increased (decreased)  $D_0$  ( $\mu$ ) for either increasing VIL or decreasing COA.  
677 On the other hand,  $N_w$  presents a different behavior where it grows with VIL up to  $VIL = 3 \text{ kg m}^{-2}$   
678 where the pattern reverses. This was consistently observed regardless of the COA values.

679 The case studies shown here provided more details on the relation between overall rain  
680 cell vertical structure and the relative COA, as well as providing a comparison between the  
681 polarimetric and DSD characteristics within the life cycle to contrast with the overall patterns  
682 found. We showed that the high COA observed at the earlier stages of the rain cells life cycles  
683 (already close to the mature stage with the thresholds used here) are associated to higher  $Z$ ,  $Z_{dr}$   
684 and  $K_{dp}$  in the mixed layer, which indicates: 1) relatively strong updrafts; 2)  $Z_{dr}$  column  
685 formation; 3) likely the beginning of cells electrification. Those characteristics were also  
686 followed by negative  $\mu$  according to the methodology adopted here. Right after this characteristic  
687 profile, there is a sudden increase in the accumulated rain in the 2 km CAPPIs. Therefore, we  
688 were able to highlight the rain cells properties that anticipate the onset of the bulk of  
689 precipitation in the lower levels. This may be used in the future as a frame of reference for new  
690 studies and to operational applications. We highlight that the COA measure does not rely on  
691 polarimetric retrievals meaning that it can provide additional insights into the rain cells  
692 properties even for single-polarization radars.

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698 <http://chuvaproject.cptec.inpe.br/portal/noticia.ultimas.logic?i=en>

699 The authors declare that there is no conflict of interests regarding this publication.  
700

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